

Relativistic correction to Hydrogen atom

①

We start with the SE: $i\hbar \frac{\partial}{\partial t} \psi = H\psi$

We can write that operator $i\hbar \frac{\partial}{\partial t}$ extracts the total energy E of the particle from the WF: $i\hbar \frac{\partial}{\partial t} \psi = E\psi$

We can write the total energy E in relativistic form:

$$E = \sqrt{c^2 p^2 + m_0^2 c^4}$$

Our purpose is to substitute operators $E = i\hbar \frac{\partial}{\partial t}$ and

Approach 1

$$E^2 = c^2 p^2 + m_0^2 c^4$$

$$-\hbar^2 \frac{\partial^2}{\partial t^2} \psi = (-c^2 \hbar^2 \nabla^2 + m_0^2 c^4) \psi$$

$$(c^2 \hbar^2 \nabla^2 - \hbar^2 \frac{\partial^2}{\partial t^2} - m_0^2 c^4) \psi = 0 \quad \text{— Klein-Gordon equation for free particle}$$

$\vec{p} = -i\hbar \vec{\nabla}$ and obtain the Schrödinger-like equation with the new Hamiltonian. To do it, we need to get rid of $\sqrt{\quad}$ in the expr. for energy.

A more general form of KG eq. will be to add electromagnetic fields:

$$E = i\hbar \frac{\partial}{\partial t} - e\Phi$$

$$\vec{p} = -i\hbar \vec{\nabla} - \frac{e}{c} \vec{A}$$

where Φ and \vec{A} are scalar and vector potentials of EM field:

$$\vec{E} = -\nabla\Phi - \frac{\partial}{\partial t} \vec{A}$$

$$\vec{B} = \nabla \times \vec{A}$$

we obtain:

$$\left[\left(i\hbar \frac{\partial}{\partial t} - e\Phi \right)^2 - c^2 \left(-i\hbar \vec{\nabla} - \frac{e}{c} \vec{A} \right)^2 - m_0^2 c^4 \right] \psi = 0$$

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For the case of H atom, we can consider EM potential as static Coulomb interaction of an e^- with the nuclei:

$$\vec{A} \equiv 0, \quad \Phi \equiv -\frac{ze}{r}$$

Then KG equation becomes:

$$\left[\left(i\hbar \frac{\partial}{\partial t} - e\Phi(\vec{x}) \right)^2 + c^2 \hbar^2 \vec{\nabla}^2 - m_0^2 c^4 \right] \psi(\vec{x}, t) = 0$$

We look for stationary solutions, so $\psi(\vec{x}, t) = \psi(\vec{x}) e^{-\frac{i}{\hbar} E t}$

$$\begin{aligned} \text{Action of the operator } \left(i\hbar \frac{\partial}{\partial t} - e\Phi(\vec{x}) \right) \psi(\vec{x}, t) &= \\ &= \psi(\vec{x}) \left(E - e\Phi(\vec{x}) \right) e^{-\frac{i}{\hbar} E t} \end{aligned}$$

$$\text{and } \left(i\hbar \frac{\partial}{\partial t} - e\Phi(\vec{x}) \right)^2 \psi(\vec{x}, t) = \psi(\vec{x}) \left(E - e\Phi(\vec{x}) \right)^2 e^{-\frac{i}{\hbar} E t}$$

So we can write the stationary KG equation:

$$\left[\left(E - e\Phi(\vec{x}) \right)^2 + c^2 \hbar^2 \vec{\nabla}^2 - m_0^2 c^4 \right] \psi(\vec{x}) = 0$$

$$\text{or } \left[\vec{\nabla}^2 + \frac{1}{c^2 \hbar^2} \left(\left(E + \frac{ze^2}{r} \right)^2 - m_0^2 c^4 \right) \right] \psi(\vec{x}) = 0,$$

This eq. has analytic solution:

$$E = m_0 c^2 \left(1 + \frac{z^2 \alpha^2}{\lambda^2} \right)^{-\frac{1}{2}}, \quad \text{where } \lambda = n + \left[\left(l + \frac{1}{2} \right)^2 - z^2 \alpha^2 \right]^{\frac{1}{2}} - \left(l + \frac{1}{2} \right)$$

Expanding in Taylor series:

$$E \approx m_0 c^2 \left(1 - \frac{z^2 \alpha^2}{2n^2} - \frac{z^4 \alpha^4}{2n^4} \left(\frac{n}{l + \frac{1}{2}} - \frac{3}{4} \right) \right)$$

relativistic correction

Approach 2 : Dirac equation

$$E = \sqrt{c^2 p^2 + m_0^2 c^4}$$

The $\sqrt{\quad}$ can be calculated if $c^2 p^2 + m_0^2 c^4 = \left(\begin{matrix} ? \\ \dots \end{matrix} \right)^2$

$$\begin{aligned} c \vec{\alpha} \cdot \vec{p} + \beta m_0 c^2 &= \\ &= c \alpha_1 p_1 + c \alpha_2 p_2 + c \alpha_3 p_3 + \beta m_0 c^2 \end{aligned}$$

We want that $\alpha_1^2 = \alpha_2^2 = \alpha_3^2 = \beta^2 = 1$
if $j \quad \alpha_i \alpha_j + \alpha_j \alpha_i = 0$
 $\forall i \quad \alpha_i \beta + \beta \alpha_i = 0$ } conditions for coefficients.

It can fulfill if $\vec{\alpha}$ and β are 4x4 matrices:

$$\vec{\alpha} = \begin{pmatrix} 0 & \vec{\sigma} \\ \vec{\sigma} & 0 \end{pmatrix}, \quad \beta = \begin{pmatrix} \mathbb{1} & 0 \\ 0 & -\mathbb{1} \end{pmatrix}, \quad \text{and} \quad \sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$$

$$\sigma_y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$$

$$\sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \text{ are}$$

Pauli matrices.

So we have $[c \vec{\alpha} \cdot \vec{p} + \beta m_0 c^2] \psi = E \psi,$

and $\psi = \begin{bmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{bmatrix}$ is a 4-vector.

We can write Dirac equation as a system of two equations:

$$(1) \quad (E - m_0 c^2) \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix} = c \vec{\sigma} \cdot \vec{p} \begin{pmatrix} \psi_3 \\ \psi_4 \end{pmatrix}$$

$$(2) \quad (E + m_0 c^2) \begin{pmatrix} \psi_3 \\ \psi_4 \end{pmatrix} = c \vec{\sigma} \cdot \vec{p} \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix}$$

We define $\psi_+ = \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix}$ solutions with $E > 0$

$\psi_- = \begin{pmatrix} \psi_3 \\ \psi_4 \end{pmatrix}$ with $E < 0$

Let us express $\psi_- = \frac{c \vec{\sigma} \cdot \vec{p}}{E + m_0 c^2} \psi_+$ from (2),

then from (1) we get

$$(E - m_0 c^2) \psi_+ = c \vec{\sigma} \cdot \vec{p} \left[\frac{1}{E + m_0 c^2} \right] c \vec{\sigma} \cdot \vec{p} \psi_+$$

Let us add EM field: ~~E~~ $F = E - e\Phi(\vec{x}, t)$

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$$\vec{P} = -i\hbar \vec{\nabla} - \frac{e}{c} \vec{A}(\vec{x}, t)$$

The Dirac equation becomes:

$$(E - e\Phi - m_0 c^2) \psi_+ = c \vec{\sigma} \cdot \vec{p} \left[\frac{1}{E - e\Phi + m_0 c^2} \right] c \vec{\sigma} \cdot \vec{p} \psi_+$$

$E = E_S + m_0 c^2$, where E_S is non-relativistic SE.

The term $\frac{1}{E - e\Phi + m_0 c^2} = \frac{1}{E_S - e\Phi + 2m_0 c^2} =$

$$= \frac{1}{2m_0 c^2} \left(\frac{1}{1 + \frac{E_S - e\Phi}{2m_0 c^2}} \right) = \frac{1}{2m_0 c^2} \left(1 + \frac{E_S - e\Phi}{2m_0 c^2} \right)^{-1}$$

We expand in series:

$$\approx \frac{1}{2m_0 c^2} \left(1 - \frac{E_S - e\Phi}{2m_0 c^2} \right) = \frac{1}{2m_0 c^2} - \frac{E_S - e\Phi}{4m_0^2 c^4}$$

We can write:

$$E_S \psi_+ = \left[\underbrace{\frac{p^2}{2m_0}}_{\text{non-relativistic part}} + e\Phi + \underbrace{c \vec{\sigma} \cdot \vec{p} \frac{E_S - e\Phi}{2m_0 c^2} c \vec{\sigma} \cdot \vec{p}}_{\text{relativistic terms}} \right] \psi_+$$

because $(c \vec{\sigma} \cdot \vec{p})^2 = c^2 p^2$

non-relativistic part

relativistic terms.

Relativistic mass correction

$$c \vec{\sigma} \cdot \vec{p} \frac{E_S - e\Phi}{2m_0 c^2} c \vec{\sigma} \cdot \vec{p} = -\frac{e\hbar^2}{2m_0 c} \vec{\sigma} \cdot \vec{\nabla} \times \vec{A} + \frac{p^4}{8m_0^3 c^2} + \frac{e\hbar^2}{8m_0^2 c^2} \vec{\nabla}^2 \Phi -$$

interaction of \vec{e} with magnetic field

$$-\frac{e\hbar}{4m_0^2 c^2} \vec{\sigma} \cdot \vec{\nabla} \Phi \times \vec{p}$$

Spin-orbit

Darwin term

Problem 1

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Proof that $\langle \psi_i | \frac{-\vec{p}^4}{8m_0^3 c^2} | \psi_i \rangle$ is equivalent to the relativistic correction in the KG equation

We use perturbation theory:

$$E_i^{(1)} = \langle \psi_i | \frac{-\vec{p}^4}{8m_0^3 c^2} | \psi_i \rangle = -\frac{1}{8m_0^3 c^2} \langle \psi_i | \vec{p}^2 \cdot \vec{p}^2 | \psi_i \rangle$$

$$H^0 | \psi_i \rangle = E_i | \psi_i \rangle$$

$$\left(\frac{\vec{p}^2}{2m_0} + V \right) | \psi_i \rangle = E_i | \psi_i \rangle$$

$$\vec{p}^2 | \psi_i \rangle = 2m_0 (E_i - V) | \psi_i \rangle$$

Thus we can write

$$E_i^{(1)} = -\frac{1}{8m_0^3 c^2} \langle \psi_i | (2m_0)^2 (E_i - V)^2 | \psi_i \rangle =$$
$$= -\frac{1}{2m_0 c^2} \left(E_i^2 - 2E_i \langle \psi_i | V | \psi_i \rangle + \langle \psi_i | V^2 | \psi_i \rangle \right)$$

For Hydrogen atom: $V(z) = \frac{-ze}{z}$

$$\langle \psi_i | V | \psi_i \rangle = \frac{-ze}{an^2}$$

$$\langle \psi_i | \frac{1}{z^2} | \psi_i \rangle = \frac{1}{a^3} \frac{1}{(l+\frac{1}{2})n^3 a^2}$$

Combining everything we get:

$$E_i^{(1)} = -\alpha^2 \frac{E_n}{n^2} \left(\frac{n}{l+\frac{1}{2}} - \frac{3}{4} \right)$$

Darwin correction:

$$\hat{D} = \frac{e\hbar^2}{8m_0^2c^2} \nabla^2 \Phi, \text{ for Coulomb potential we have}$$

$$\Phi = -\frac{ze}{r}$$

$$\hat{D} = \frac{\pi\hbar^2 ze^2}{2m_0^2c^2} \delta^3(\vec{r})$$

$$\langle \hat{D} \rangle = \frac{\hbar^2}{8m_0^2c^2} 4\pi \left(\frac{ze^2}{r}\right) |\psi(0)|^2$$

$$\psi(0) = 0 \text{ for } l > 0$$

$$\psi(0) = \frac{1}{\sqrt{4\pi}} 2 \left(\frac{z}{na_0}\right)^{3/2} \text{ for } l = 0$$

$$\langle \hat{D} \rangle = \frac{2\pi}{m_0c^2} E_n^2, \text{ so the Darwin term affects only the s orbitals.}$$

Spin-orbit term:

$$-\frac{e\hbar}{4m_0^2c^2} \vec{\sigma} \cdot \nabla \Phi \times \vec{p} \text{ for Coulomb potential}$$

$$\Phi = -\frac{ze_0}{r}$$

$$\nabla \Phi = -\frac{ze_0}{r^3} \vec{r}$$

$$\frac{e}{2m_0^2c^2} \frac{\hbar^2}{2} \vec{\sigma} \cdot \frac{ze\vec{r}}{r^3} \times \vec{p} = \frac{ze^2}{2m_0^2c^2 r^3} \vec{\sigma} \cdot \vec{L}, \text{ because } \vec{r} \times \vec{p} = \vec{L}$$

The spin orbit interaction = 0 for s states because $\vec{L} = 0$.

In other cases we need to find correction proportional $\sim \langle \psi | \vec{S} \cdot \vec{L} | \psi \rangle$.

Spin orbit coupling:

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We work with the Hamiltonian in the form:

$$\hat{H} = \underbrace{-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{x})}_{H_0} + \underbrace{\frac{ze^2}{2m^2 c^2 \hbar^3} \vec{S} \cdot \vec{L}}_{\Delta H}$$

We know the eigenstates and energies of the Hamiltonian \hat{H}_0 .
In principle, we can use perturbation theory to find E of \hat{H} .

$$\hat{H}_0 |n^{(0)}\rangle = E_n^{(0)} |n^{(0)}\rangle$$

First order correction:

$$E_n^{(1)} = \langle n^{(0)} | \Delta H | n^{(0)} \rangle = \frac{ze^2}{2m^2 c^2} \langle n^{(0)} | \frac{1}{\hbar^3} \vec{S} \cdot \vec{L} | n^{(0)} \rangle$$

Problem is to act by operator $\vec{S} \cdot \vec{L}$ on unperturbed states $|n^{(0)}\rangle$

What the operator $\vec{S} \cdot \vec{L}$ means?

$$\vec{S} \cdot \vec{L} = \hat{S}_1 \otimes \hat{L}_1 + \hat{S}_2 \otimes \hat{L}_2 + \hat{S}_3 \otimes \hat{L}_3 = \sum_i \hat{S}_i \otimes \hat{L}_i$$

We can use the summation of angular momenta:

$$\vec{J} = \vec{L} + \vec{S} \equiv \vec{L} \otimes \mathbb{1} + \mathbb{1} \otimes \vec{S}$$

$$\vec{J}^2 = \sum_i \hat{J}_i \cdot \hat{J}_i = \vec{L}^2 \otimes \mathbb{1} + \mathbb{1} \otimes \vec{S}^2 + 2 \underbrace{\sum_i \hat{S}_i \otimes \hat{L}_i}_{\vec{S} \cdot \vec{L}}$$

$$\vec{S} \cdot \vec{L} = \frac{1}{2} (\vec{J}^2 - \vec{L}^2 - \vec{S}^2)$$

$$[\hat{J}^2, \hat{L}^2] = 0$$

$$[\hat{J}^2, \hat{S}^2] = 0$$

It can be shown that

$$[\hat{L}^2, \hat{L}_i] = 0 \quad [\hat{L}_i, \hat{S}_i] = 0$$

$$[\hat{S}_i, \hat{S}_i] = 0 \quad [\hat{L}^2, \hat{S}_i] = 0$$

$$[\hat{S}^2, \hat{L}_i] = 0$$

Hydrogen atom:

Commuting observables: $\{H, L^2, L_z\}$, for \hat{H}_0

which we can expand to add $\{S^2, S_z\}$

What is the set of comm. observables for

$$\hat{H} = \hat{H}_0 + \Delta \hat{H} \quad ? \quad \{H, L^2, S^2, J^2, J_z\},$$

$\left. \begin{matrix} \vec{S} \cdot \vec{L} \\ \cdot \end{matrix} \right\}$

but L_z and S_z do not commute with H anymore.

The original state $|l, m\rangle$ and $|S, m_s\rangle$ form new basis by tensor product: $|l, m\rangle \otimes |S, m_s\rangle$.

Consider the case $n=2, l=1$ (2p orbital of H):

We have

$ 1, 1\rangle$	$ 1/2, 1/2\rangle$	
$ 1, 0\rangle \otimes$	$ 1/2, 1/2\rangle$	and will obtain in total
$ 1, -1\rangle -$	$ 1/2, -1/2\rangle$	6 new states
(3 states)	(2 states)	

Let us classify the states according to $\hat{J}_z = \hat{L}_z + \hat{S}_z$:

$$\begin{aligned}
 +\frac{3}{2} &: |1, 1\rangle \otimes |1/2, 1/2\rangle \\
 +\frac{1}{2} &: |1, 0\rangle \otimes |1/2, 1/2\rangle, \quad |1, 1\rangle \otimes |1/2, -1/2\rangle \\
 -\frac{1}{2} &: |1, 0\rangle \otimes |1/2, -1/2\rangle, \quad |1, -1\rangle \otimes |1/2, 1/2\rangle \\
 -\frac{3}{2} &: |1, -1\rangle \otimes |1/2, -1/2\rangle
 \end{aligned}$$

We need to form states which are eigenstates of \hat{J}^2 (9)

We showed that for $l=1, s=\frac{1}{2}$, we can form $J_z = \left\{ +\frac{3}{2}, +\frac{1}{2}, -\frac{1}{2}, -\frac{3}{2} \right\}$

Thus we obtain possible $J = \left\{ \frac{3}{2}, \frac{1}{2} \right\}$

\uparrow \uparrow
 4 states 2 states

$\left\{ +\frac{3}{2}, +\frac{1}{2}, -\frac{1}{2}, -\frac{3}{2} \right\}$ $\left\{ +\frac{1}{2}, -\frac{1}{2} \right\}$

Now we form the full set:

(1) $|J = \frac{3}{2}, m = \frac{3}{2}\rangle = |1, 1\rangle \otimes |\frac{1}{2}, \frac{1}{2}\rangle$ - only one state satisfies this condition.

Same is for

$$|J = \frac{3}{2}, m = -\frac{3}{2}\rangle = |1, -1\rangle \otimes |\frac{1}{2}, -\frac{1}{2}\rangle$$

for $|J = \frac{3}{2}, m = \frac{1}{2}\rangle$ it should be a linear combination of $|1, 0\rangle \otimes |\frac{1}{2}, \frac{1}{2}\rangle$ and $|1, 1\rangle \otimes |\frac{1}{2}, -\frac{1}{2}\rangle$

We can obtain the coefficients by using raising and lowering operators:

$$J_{\pm} |J, m\rangle = \hbar \sqrt{J(J+1) - m(m \pm 1)} |J, m \pm 1\rangle$$

Applying ~~$J_{\pm} |J = \frac{3}{2}, m = \frac{3}{2}\rangle$~~

$$J_{-} |J = \frac{3}{2}, m = \frac{3}{2}\rangle = \hbar \sqrt{\frac{3}{2} \left(\frac{5}{2}\right) - \frac{3}{2} \left(\frac{1}{2}\right)} |J = \frac{3}{2}, m = \frac{1}{2}\rangle = \hbar \sqrt{3} |J = \frac{3}{2}, m = \frac{1}{2}\rangle$$

We apply it now for the RHS^{of (1)}, so

$$\begin{aligned}
 J_{-} |1, 1\rangle \otimes |\frac{1}{2}, \frac{1}{2}\rangle &= (L_{-} + S_{-}) |1, 1\rangle \otimes |\frac{1}{2}, \frac{1}{2}\rangle = \\
 &= \hbar \sqrt{2} |1, 0\rangle \otimes |\frac{1}{2}, \frac{1}{2}\rangle + \hbar |1, 1\rangle \otimes |\frac{1}{2}, -\frac{1}{2}\rangle
 \end{aligned}$$

Finally we get

$$|J = \frac{3}{2}, m = \frac{1}{2}\rangle = \sqrt{\frac{2}{3}} |1, 0\rangle \otimes |\frac{1}{2}, \frac{1}{2}\rangle + \frac{1}{\sqrt{3}} |1, 1\rangle \otimes |\frac{1}{2}, -\frac{1}{2}\rangle$$

Similarly, we get

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$$|\frac{3}{2}, -\frac{1}{2}\rangle = \sqrt{\frac{2}{3}} |1, 0\rangle \otimes |\frac{1}{2}, -\frac{1}{2}\rangle + \frac{1}{\sqrt{3}} |1, -1\rangle \otimes |\frac{1}{2}, \frac{1}{2}\rangle$$

and

$$J = \frac{1}{2} \quad |\frac{1}{2}, \frac{1}{2}\rangle = -\frac{1}{\sqrt{3}} |1, 0\rangle \otimes |\frac{1}{2}, \frac{1}{2}\rangle + \sqrt{\frac{2}{3}} |1, 1\rangle \otimes |\frac{1}{2}, -\frac{1}{2}\rangle$$

$$|\frac{1}{2}, -\frac{1}{2}\rangle = \frac{1}{\sqrt{3}} |1, 0\rangle \otimes |\frac{1}{2}, -\frac{1}{2}\rangle + \sqrt{\frac{2}{3}} |1, -1\rangle \otimes |\frac{1}{2}, \frac{1}{2}\rangle$$

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More generally:

$$|nl s j m_j\rangle = \sum_{m_l, m_s} \underbrace{\langle l s m_l m_s | j m_j \rangle}_{\text{Clebsch-Gordan coefficients}} |nl s m_l m_s\rangle$$

Finally, we have

$$\langle \vec{L} \cdot \vec{S} \rangle = \langle \frac{1}{2} (J^2 - L^2 - S^2) \rangle = \frac{\hbar^2}{2} (J(J+1) - l(l+1) - s(s+1))$$

Which is for the case of $l=1, s=\frac{1}{2}$ is

$$= \frac{\hbar^2}{2} (J(J+1) - 2 - \frac{3}{4}) = \frac{\hbar^2}{2} (J(J+1) - \frac{7}{4})$$

